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The coupling between atomic and electronic structure in small Cu clusters

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Abstract. Thermodynamic and ground-state properties of Cu clusters have been studied with **the** effective-mediun theory including a tight-binding descriptian of **the** one-electron speclnun. Simulated-annealing Monte carlo calculations have **been** perfmed for cluster sizes between 3 and 29 to determine ground-state energies and structures. Finite-temperature ensembles have been generated for a range of temperatures. The magic numbers 8, 18 and 20 are reproduced and remain stable at temperatures close to 1000 K, where the clusters may be regarded as liquid. The coupling between the atomic and electronic degrees of freedom through a Jahn-Teller-like effect is shown to play a key role in understanding the stability of the magic clusters *at* **high** tempera". For the nan-mapic clusters this Jahn-Teller effect gives **rise** to large Fermi **gaps** even at high temperatures and a pronounced stability of the even-sized clusters relative to *those* with odd sizes. Finite-temperature electronic spectra are calculated. The fluctuations in the atomic positions give rise to a large broadening of the electronic levels. in agreement with experimental observations. Cu₁₃ exhibits a rather sharp melting transition, whereas clusters of other sizes show more complex behaviour.

1. Introduction

The physics *of* metal clusters **is** a field that has developed rapidly over the last few years [l]. Clusters of certain sizes (magic numbers) exhibit an extraordinary stability [I], in much the same way **as** is the case with nuclei. Magic numbers are seen in abundance measurements, ionization potentials, dissociation energies etc. They were first observed for Na [21, but the same magic numbers have also been **seen for** other monovalent free-electron-like metals. Katakuse and co-workers 131 performed abundance measurements for positively charged clusters of Cu, Ag and **Au.**

The magic numbers correspond **to** closed angular momentum shells, and this idea has been illustrated in a number of simple models with the common characteristic that the ion cores have only been taken into account in a very approximate way. Clemenger **[41** applied the Nilsson model [5] *of* nuclear physics to metallic clusters. Here, the one-electron levels are constructed **as** eigenvalues of a simple spheroidal harmonic oscillator. In a study by Ekardt **[61** the ion cores were replaced by a homogeneous positively charged sphere. Localdensity calculations on the system *of* valence electrons reproduced the magic **numbers** observed **for** Na. In this model the geometry *of* the cluster was initially ignored Later, spatial structure was introduced through a spheroidal distortion of the jellium spheres [7].

Calculations including the atomic cores have been performed at different levels *of* approximation. Cluster sizes larger than 13 have been treated almost entirely by very simple

models. **An** extensive calculation with the simple Hiickel model [SI has been performed by Lindsay and co-workers. Small clusters have been treated in configuration-interaction calculations; for instance, the relaxed ground-state of Na-clusters of up to eight atoms have been determined **[9].** Local-density calculations including the ion-cores have been done mainly for the ground state of small alkali clusters consisting of a maximum of 13 atoms $[10, 11]$. Andreoni and co-workers $[12]$ also performed calculations for Na₁₈ and $Na₂₀$ with the Car-Parrinello method, where the ground state was found with simulated annealing and where finite-temperature properties were investigated.

The experimental knowledge of the atomic structure of clusters is very sparse. Only the structures of some clusters up to seven atoms have been determined [13].

In the present work we have applied a model for the binding in a metal cluster, which includes a description of the atomic degrees of freedom but which is still computationally sufficiently simple to allow for studies of many cluster sizes over a wide temperature range. In the following section we will describe the model, the effective-medium theory [14] including a tight-binding description of the one-electron spectrum. In the third section we describe and discuss the results of applying the model to small **Cu** clusters. We shall discuss the interplay between 'magicity' and geometry, and show that atomic relaxations are very important in understanding the extraordinary stability of clusters with magic sizes. Sizeable Fermi gaps also exist in non-magic clusters with **an** even number of atoms. The influence of heating of the atomic degrees of freedom on stability is treated, and the effect of high temperatures on the electronic spectrum is discussed. The melting of small clusters is also studied. It is shown that Cu₁₃ has a well defined melting transition at $T \simeq 450$ K. Other cluster sizes exhibit a more complex behaviour.

A brief account of some of the results discussed here has been presented in [15].

2. The effective-medium model

To study the coupling between the atomic and electronic degrees of freedom in small **Cu** clusters we construct a model which can be used to calculate the total energy and the electronic structure of a cluster **as a** function of atomic positions. The method applied in the present work is the effective-medium theory (EMT) [14]. It has the property of reproducing a well tested interatomic potential for an infinite system of metallic Cu and it takes finite-size shell effects into account through a tight-binding description. The effective medium model is based on density functional theory and relies on the fact that errors in the total electron density give rise to only second-order errors in the total energy due to the variational properties of the energy functional. The idea behind the **EMT** is to approximate the energetics of an atom in a metallic environment by a simple reference system, which can be taken **as** the atom embedded in a homogeneous electron gas or **as** the atom situated in a perfect **FCC** crystal with a suitable lattice constant. In the following presentation we shall use the **FCC** crystal **as** the reference system. The **Em** total-energy expression may be written **as**

$$
E_{\text{tot}} = \sum_{\text{atoms } i} E_{\text{c}}(\bar{n}_i) + \sum_{\text{atoms } i} E_{\text{AS}}^{(i)} + E_{1-\text{el}}
$$
(1)

as described in detail elsewhere $[14]$. The cohesive energy function E_c describes the energy of an **FCC** crystal of a lattice constant corresponding to a background electron density of *ii* felt by each atom. The atomic-sphere correction energy *EAS* represents the repulsion between overlapping atomic spheres and plays an important role in non-close-packed structures where

the overlap is significant. The one-electron correction $E_{1-\text{el}}$ is a correction concerning the difference **in** density of states **(DOS)** between the reference system and the real system.

For a finite system like a cluster of Cu atoms the electron **DOS** consists of a discrete set of levels which are not well represented by the **DOS** of an infinite FcC Cu crystal. It is therefore necessary to include the one-electron energy in the description of the energetics of small clusters. The one-electron energy $E_{1-\epsilon i}$ is the energy difference due to the formation of electronic shells instead *of* a continuum **as** in the extended FCC system, and we shall therefore in the following use the term *shell energy* to be synonymous with the one-electron energy.

The important shell structure in small Cu clusters is due to the **4s** electrons. The centre of the 4p band of bulk **Cu** is situated approximately one Rydberg above the centre *of* the 4s band **[16],** and we shall therefore ignore atomic p orbitals in **our** description. We shall also ignore the filled 3d orbitals. In a more accurate description the **p** and d orbitals could be included; but since the one-electron energy is only a small correction term which describes deviations in the electronic structure from that in **an** FCC crystal, a simple model of the **^s** band is sufficient.

The bonding between the atomic **s** orbitals is described with a tight-binding Hamiltonian. We assume **for** simplicity that all the atomic **s** levels **are at** the same energy and choose our energy zero at this common energy. We determine the distance dependence of the off-diagonal matrix elements *Hij* between atomic **s** orbitals situated at atoms *i* and *j* in the following way. The matrix elements can be expected to decay in the **same** way **as** an orbital on one of the atoms, which again falls off roughly **as** the square root of the electron density. In the construction **of** the embedding densities an exponential fall-off *is* used

$$
\bar{n}_i \propto \sum_{j \neq i} e^{-\eta_2 r_{ij}} \tag{2}
$$

where η_2 is a parameter characteristic of Cu_t. We therefore take the hopping matrix elements to be of the form

rm
\n
$$
H_{ij} = -V_0 e^{-(\eta_2/2)(r_{ij}-r_0)} \t i \neq j
$$
\n(3)

where r_0 is the equilibrium nearest-neighbour distance between Cu atoms in an FCC crystal. The prefactor V_0 is determined so that the width of the occupied part of the s band in the tight-biding model for an FCC crystal **at** equilibrium volume agrees **with** the one calculated self-consistently by Moruzzi and co-workers **[17].** The value obtained in this way is $V_0 = 0.567 \text{ eV}$.

For a given atomic configuration of a cluster the tight-binding Hamiltonian is used to calculate the discrete electron spectrum with energies ϵ_{α} of the cluster. The band energy *E*_{band} in the reference FCC system should be subtracted in the one-electron energy correction. The band energy is a function of the embedding density \bar{n} , and this function can be calculated once and for all by solving the tight-binding Hamiltonian for a range of lattice constants of an FCC lattice. The shell energy therefore takes the form

$$
E_{1-\text{el}} = \sum_{\text{occupied}} \epsilon_{\alpha} - \sum_{i} E_{\text{band}}^{\text{FCC}}(\bar{n}_{i}). \tag{4}
$$

To summarize, we have described an effective-medium model for the total energy of a cluster as a function of atomic positions. The model consists of energy terms which describe the local volume changes through embedding densities, electrostatic interactions by use **of** pair potentials, and the shell energy by a tight-binding Hamiltonian.

t **A** table with all **the** effective-medium parameters for Cu *can* be **found in [14].**

3. Results and **discussinn**

We have performed a number of simulated-annealing Monte Carlo calculations with the Metropolis algorithm [181 to determine the ground-state configurations of small Cu clusters. The simulations have been done both with and without the inclusion of the shell energy E_{1-1} . Typically 2-5 runs of 2000 coordinate sweeps were made at a given cluster size, starting at temperatures in the range 2000–10 000 K. A sweep involves one trial movement of each coordinate, i.e. *3N* total-energy calculations. Finite-temperature ensembles were generated using 10000-40000 sweeps.

3.1. Ground-state properties

In this section we present and discuss the calculated ground-state properties of the clusters. First, we present the obtained ground-state energies and atomic structures and compare them with results from other models and experiment in the cases where it is possible. We then discuss the coupling between the **electronic** and atomic structure via the Jahn-Teller effect. This coupling is crucial in understanding the extraordinary stability of the magic clusters and the pronounced even-odd oscillations observed in experimental abundance spectra **[31.**

The calculated ground-state energies are depicted in figure 1. We measure E_c from the minimum value, which is 3.56eV below the atomic energy. Through the points a smooth three-parameter curve is drawn with the form

$$
E_{\text{fit}} = aN + bN^{2/3} + cN^{1/3} \tag{5}
$$

where the first term represents a **bulk** energy per atom, the second represents a surface energy term and the third a curvature term. The magic cluster sizes show up as tiny oscillations on the curve. As a convenient way of extracting the magic numbers we plot in figure 2 the deviation of the total energies from the fit (5).

Figure 1. Toral **energy as a function** *of* **cluster size. The triangles are obtained with a full calculation while the squares result from a calculation where** the **shell** energy $E_{1-**el**}$ is neglected. The curves are fits to the **data with the analytic** form of **(5).**

Figure 2. Deviation of total energy from the fit (5). **The triangles are obtained with a full** calculation while **the squares result from a calculation where the shell** energy E_{1-e} is neglected.

A different way of enhancing the fine **structure** is to take the second difference

$$
\Delta_2(N) = E(N+1) + E(N-1) - 2E(N). \tag{6}
$$

Figure 3. Second difference of the total energy, Figure 4. Calculated ground-state structures for $N = 5$, **Figure 3.** Second difference of the total energy, $E(N + 1) + E(N - 1) - 2E(N)$. The triangles are **obliiined with a** full **calculatron while the squares result** from a calculation where the shell energy $E_{1-**el**}$ is **neglected.**

7. **10 and 13.**

This quantity is shown in figure **3.** Assuming thermal equilibrium it follows [l] that the value of $\Delta_2(N)$ is proportional to the logarithm of the abundance of N clusters in an experimental **beam.**

It is clear from the figures that the clusters of size 8, 18 and 20 are particularly stable. This means that the tight-binding form used to model the shell energy is indeed able to reproduce the magic numbers seen in experiment **[Z,** 31 and in other models **[2,4,6].** The magic numbers do not appear in the calculations where the shell energy is neglected. Without the shell energy the ground-state configurations are close-packed and the sizes 13 and 19 appear **as** magic corresponding to close-packed ground-state. structures with a central atom surrounded by all atoms up to first and second nearest-neighbour shells, respectively.

The s-state tight-binding model for calculating the shell energy is general for monovalent metals, and the ground-state structures found in the present calculation *are* in fact very similar to previously published theoretical [8-10, 19] and experimental [13] results for alkali clusters. For instance, in the case of $N = 7$ the equilibrium structure is a fivefold bi-pyramid (see figure 4) as indicated by electron spin resonance experiments on Li [13]. For $N = 13$ we find a cuboctahedral rather than an icosahedral geometry like the Na result **of** [lo]. Structures for cluster sizes $N = 5, 7, 10$ and 13 are shown in figure 4. In figure 5 we compare the structure for the $N = 20$ Cu cluster with the structure for $Na₂₀$ found with the Car-Paninello method 1121 by Ballone and co-workers. The two structures **are** very similar, indicating that the cluster structures for the two different metals are determined by the same mechanisms.

To study the variation of the shape of the clusters **as** a function of size we have evaluated the eccentricities. We define the eccentricity of a cluster as follows. The inertial (3×3) matrix is defined **as** usual as

$$
\overline{\overline{J}}_{ij} = \sum_{k} [|\boldsymbol{r}_k - \tilde{\boldsymbol{r}}|^2 \delta_{ij} - (\boldsymbol{r}_k - \tilde{\boldsymbol{r}})_i (\boldsymbol{r}_k - \tilde{\boldsymbol{r}})_j] \tag{7}
$$

where *i*, $j = 1, 2, 3$ are the coordinates and $\bar{r} = (1/N) \sum_{k} r_k$ is the centre of mass of the cluster. The three eigenvalues *J,* of this matrix are the moments of inertia. In the case of an ellipsoid of axes a_i , these are also principal inertial axes with $J_i \propto a_i^2$. Thus we define the two eccentricities of a cluster **as**

$$
\epsilon_{1,2} = \sqrt{J_{1,2}/J_3} \qquad J_1 \leqslant J_2 \leqslant J_3. \tag{8}
$$

r

i

,, **Figure 5.** Ground-state structure for $N = 20$ from our calculation (top **panel)** and from the **local-density** calculation of [12] for a Na₂₀ cluster (bottom panel).

, ,,. ,, , ,

,, , , , ,, ,, ,, , , , , , , , ,

,

In figure 6 we plot the two calculated eccentricities for the ground-state structures. For comparison we also show eccentricities determined from a calculation by Selby **and** coworkers 1201. In the latter calculation the principal axes are related to vibrational frequencies of a three-dimensional harmonic oscillator potential, **and** the energy is minimized with respect to one distortion parameter. This so-called Nilsson model **[SI** was first applied to metallic clusters by Clemenger **[4].**

Figure 6. Eccentricities of ground-state structures from the present model (top **panel) and from the** Clemenger **model** [241 **(bottom panel).**

Figure 7. One-electron levels in the calculated cluster **gound-state configurations.** The **highest occupied levels are indicated by dots** (one **for each electron on the twel).**

As predicted by the Clemenger model, the magic clusters prefer a spherical geometry, whereas non-magic clusters tend to be distorted. One feature present in **our** calculation, but not in the Clemenger model, is the almost spherical geometry of $N = 13$. The minimization of the electrostatic repulsion **of** the ionic nuclei leads to a close-packed cuboctahedral structure of the 13 cluster. This fact was already established in local-density calculations by Martins and co-workers **[IO].** The observation that the tendency to close-packing may

be important for determining the ground-state structure, at the same time **as** the total energy has a size-dependence corresponding to angular momentum level bunchings, has been established experimentally for larger Cu clusters [21].

We now tum to a more detailed discussion of the electronic structure which results from the present model and the interplay with the atomic arrangement.

The tight-binding Hamiltonian **(3)** has only *N* eigenvalues and it is therefore clearly unable to reproduce the full infinite spectrum of angular momentum levels. However, as has been shown for instance in the framework of the LMTO method [I61 a tight-binding description may give a good representation of the electronic stmcture in a **certain** energy window. In figure 7 we plot the one-electron spectrum of the equilibrium structure **as** a function **of** cluster size. The lower levels bunch up in a lowest band of one state and a second-lowest of three states, thus reproducing one **s** and three p states. For the largest clusters a d band of five states is recognized. The grouping of levels in **s,** p and d shells is also seen in the spectra calculated by Lindsay and co-workers [SI **for** the gmund states of simple Hiickel clusters. In both calculations, though, the similarity to an angular momentum spectrum is only valid some levels below the Fermi level at *N/2.*

At all even cluster sizes, the ground state has a noticeable Fermi gap irrespective of whether there should be a shell closing or not according to the angular momentum model. This is illustrated in figure 8. The gap is largest at the magic numbers, but is quite large at all sizes. Also shown in figure 8 are experimentally observed gaps for $Cu_N⁻$ determined with photoelectron spectroscopy [22]. We see that the calculated gaps **are** in fair agreement with the experimentally determined values both with regard to the absolute size and the trend when the size of the cluster is changed. Note in particular that the ratio between a typical gap for a magic and a non-magic cluster is well reproduced. This ratio is clearly exaggerated in simpler models where the atomic degrees *of* freedom are neglected [22]. Oscillations in the cluster energy are seen in experiment [1,3], but **are** *nor* of any appreciable size in the jellium-ball model of Ekardt *[6]* or the Clemenger model **[4]** compared with the strength of the magic numbers. One should bear in mind that the experimentally measured gap applies to geometries assumed by negatively charged clusters and is therefore different from the gap we calculate for the ground state *of* the neutral species-it does, however, give an estimate of the relevant order of magnitude of the quantities considered.

The existence of Fermi gaps for the non-magic clusters and the observation of an evenodd oscillation of the ground-state energy as a function of cluster size lead to the conjecture that a Jahn-Teller-like effect is decisive **in** determining the ground-state structures **[23].** *In* the case of an even-sized cluster a small distortion can give rise to the opening of a gap between the occupied and the unoccupied levels with a gain in the total energy as a result.

To confirm that the Jahn-Teller effect is responsible for the even-odd oscillation, we have calculated the electronic spectra and the total energies *of* clusters in two alternative sets of atomic structures. The first of these, which we shall call the continuum structures, are the ground-state structures for the energy expression (1) without the shell energy $E_{1-\text{el}}$. These structures represent ground-state structures if the interatomic interactions **are** like in the bulk metallic system where the electronic spectrum is continuous. The second set of structures are simply pieces of an **FCC** crystal which **are** chosen to **be** as spherical as possible. This set we call the **FCC** structures. The lattice constant for the **FCC structures are** chosen to minimize the full energy expression (1). **In** figure 9 we show one-electron levels for clusters of size $N = 16$ and $N = 20$ in the three different structures: the ground-state structure, the continuum structure, and the **FCC** structure. It is clear that the gap is much smaller for the latter two atomic structures. **Also,** if we consider the total energy evaluated with the full expression (1) versus cluster size for the continuum or **FCC** structures (figure 10) we see no special stability of the magic sizes *8,* 18 and 20. This **lack** of extra stability at the magic sizes is not due to the overall shape of the clusters because the continuum and **FCC** structures are **just as** spherical as the ground-state structures. This is therefore a strong indication that the relaxation of the atomic structure is of crucial importance in explaining the stability of the magic clusters. The atomic and electronic structure must therefore be treated simultaneously as *in* the present model so the atomic positions are allowed to relax to minimize the total energy including the shell energy.

Figure **10. Second energy difference evaluated using (1).** The three structures are **(A)** the ground-state structures, (B) the continuum structures found by **neglecting the shell energy. and (C) a maximally spherical section of an** FCC **lattice. Only the sequence of fully relaxed ground-state svuctures show the magic numbers** *8,* **18 and 20. The even-odd fluctuation of the** $energy$ is also absent in the *two* lower curves.

Figure 11. Energy deviation from the *fit* **(5). Top panel:** $T = 0$. Middle panel: $T = 460$ K. Bottom **panel:** $T = 930$ K. The magic numbers 8, 18 and 20 **remain stable.**

In summary, we have seen that the effective-medium model reproduces many qualitative ground-state features of simpler jellium and Huckel models and also agrees with **more** firmly based first-principles calculations and experimental results where available. The model indicates that the coupling between the atomic and electronic structures via relaxations is of importance in explaining the extraordinary stability of large clusters at the magic numbers and also in understanding the electronic spectra, which also exhibit sizable Fermi gaps **for** clusters **of** (even) non-magic sizes, **as** observed with photoelectron spectroscopy 1221.

3.2. Finite-temperature properties

In this section we discuss some aspects of the behaviour of small **Cu** clusters at high temperatures. We shall show that the extraordinary stability of the magic clusters survive even in a temperature regime where the clusters can be regarded **as** melted. We also discuss the electronic spectrum at elevated temperatures. Due to the coupling of the electronic structure to the atomic degrees of freedom the spectral peaks are broadened considerably.
The peak widths are proportional to \sqrt{T} . ~.

In figure **I1** we plot the deviation from the fit (5) **(as** in figure *2)* for Monte Carlo generated ensembles at $T = 460 \text{ K}$ and $T = 930 \text{ K}$. The energy variation at $T = 460 \text{ K}$ is almost identical to that in the ground state, and even though the minimum at $N = 20$ is somewhat weaker at $T = 930$ K'than at $T = 0$, the spectra are very similar. The magic numbers 8, 18 and 20 prevail even at this high temperature;

To analyse the high-temperature configurations we plot in figure 12 the thermal average **of** the one-electron spectrum of the non-magic 16 cluster and the cluster at the magic size $N = 20$, both at the temperature $T = 930$ K. For both the 16 cluster and the 20 cluster we observe a considerable broadening of the electronic levels due to the atomic motion, but an appreciable Fermi gap is still observed for both sizes. The broadening of the levels are seen to be several times $k_B T$. Similar level broadening is seen in experiment [22]. This is a consequence of the fact that only the total energy, and not the one-electron levels, are at a minimum in the ground-state structure of a cluster. Therefore there will be first-order variations in the position of a level **as** a function of coordinate fluctuations. In the harmonic approximation the size of coordinate fluctuations is proportional to \sqrt{T} , be proportional to \sqrt{T} , as opposed to the total energy, which is proportional to *T*. Even though the one-electron peak widths are large, the total energy only varies with thermal energies $\sim \sqrt{(3N-6)/2} k_B T$. For this reason it is still possible to see magic numbers at elevated temperature. and therefore the leading term in the temperature variation of a one-electron level will also

The clusters at $T = 930$ K may be regarded as liquid in many respects, as we shall now show. The bulk melting point of copper is 1356 K, but it is well known that surfaces can introduce disorder or pre-melting at temperatures below the bulk melting temperature *[XI.* This phenomenon has also been studied using simulations based on the effective-medium theory *[U].* The clusters are finite systems, and therefore we do not observe a phase transition in the usual thermodynamic sense, and there may even exist a temperature regime in which solid- and liquid-like phases coexist [26]. In the following we shall present some calculated finite-temperature values for quantities that give indications of the temperature at which the clusters can be regarded **as** melted: the eccentricity and the radial and angular correlation functions.

As a signature of the fluctuations in the overall shape of a cluster at a particular temperature we plot the average value of the smallest eccentricity ϵ_1 (8) in figure 13 for all cluster sizes at the same temperatures as in figure 11. We first observe that at $T = 460 \text{ K}$ the spherical nature of the 13 cluster has been suppressed so that it falls more naturally on the Clemenger curve for the eccentricities. This is an indication that the shell energy is important at this temperature compared with the bulk-metal-like interactions that drove the 13 cluster spherical in the ground state. Most of the pronounced peaks in the curve have

 $T = 0$
 $T = 0$
 $T = 460K$
 $T = 930K$
 $T = 930K$

Cluster size

1.3. Average smallest eccentricity ϵ_1 . Full
 $T = 0$. Broken curve: $T = 460K$. Chain
 $T = 930K$. The structure of the spectrum is
 $T = 930K$. The structure E ceentricity
 0.8
 0.6 **0.4 0** *5* **10 15 20** *25* 30 ciusier **sire**

 10

Figure 12. Temperature-averaged one-electron **spec**trum for the cluster $N = 20$ at $T = 930$ K. Broken curve: average E_F . Full vertical lines: energy levels from equilibrium structure, i.e. the one-electron spec**uum at zero** temperature. **A** large dip in Ihe *ws at* the Fermi level is still clearly seen even at this high temperature.

Figure 13. Average smallest eccentricity ϵ_1 . Full curve: $T = 0$. Broken curve: $T = 460$ K. Chain curve: $T = 930$ K. The structure of the spectrum is smeared **out** at the higher temperaturs; at the highest temperature the magic cluster $N = 20$ is no longer more spherical than its neighbours.

disappeared at $T = 930$ K, which is a sign that the clusters at this temperature sample a much larger region of phase space. However, **as** noted above, the clusters **maintain** appreciable Fermi gaps even at **this** temperature due to local Jahn-Teller distortions.

The detailed atomic structure is probed by the radial and angular correlation functions. In figure **14** the angular correlation function is shown for the 20 cluster for a range of temperatures. We observe that the pronounced *peaks* in the low-temperature **spectra** for the second, third and fourth nearest-neighbours disappear in a temperature regime around 230460K. This indicates a transition in the cluster *to* **a** more disordered 'phase' at this temperature.

Figure **14.** Angular correlation function at cluster size $N = 20$. A cutoff radius of 5.60 bohr was used, cf the previous figure. From bottom to top: $T = 0$ ground **state.** $T = 230 \text{ K}$, $T = 460 \text{ K}$ and $T = 930 \text{ K}$ thermal averages. It **is** clearly indicated that the smcture melts between the two temperatures $T = 230$ K and $T = 460$ K.

Figure 15. Heat capacity of Cu₁₃ and Cu₂₀ calculated [27] from energy distributions of Monte Carlo **runs** at different *T.* There **is** one Well defined feature at $T \simeq 450$ K for Cu₁₃, indicating a finite-size smeared phase transition. The 20 cluster exhibits a complex behaviour with a deviation from the harmonic behaviour $c_p = (3N - 6)/2$ setting in below $T \approx 200$ K.

On the basis of energy distributions for *a* whole set of Monte Carlo runs at different *T* it is possible [27] to give estimates of the classical density of states $W(E)$ such that

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thermodynamic averages of functions of energy $f(E)$ can be calculated as

$$
\langle f \rangle_T = \int dE \, f(E) \, W(E) \, e^{-E/k_B T}.\tag{9}
$$

Using this method we have calculated the heat capacity c_p as a function of *T* for different cluster sizes. In figure 15 we show the results for $N = 13$ and $N = 20$. For Cu₁₃ there is a clear indication of a smeared-out phase transition at $T \simeq 450$ K, whereas the picture for Cu₂₀ is difficult to interpret. Studies of Lennard-Jones clusters [28] have indicated clear melting transitions for cluster sizes $N = 13$, 55 and 147, corresponding to closed Mackay icosahedra. Guided by the above-mentioned results for eccentricity and correlation functions as well as by observations of the configurations assumed by Cu_{13} at different temperatures it is possible to correlate the feature in the c_p curve with a transition from configurations with one well defined central atom to open configurations with no highly coordinated atoms.

The rise of c_p for Cu₂₀ at very low T may be associated with soft shape oscillations or the possible existence of different isomers, **as** is found to be the case in recent theoretical studies of small Na and Si clusters **1291.** The underlying mechanisms have not yet been sorted out in any detail. The results seem to indicate that cluster sizes corresponding to close-packed configurations are not necessarily representative of the thermodynamic properties of small clusters, and that further investigations of other sizes are necessary. This is also concluded in a recent paper [30] on small Be clusters. In this work, Be₁₃ is seen to have a higher melting point than smaller Be clusters, probably as a consequence of the close-packed ground-state structure.

4. Conclusion

The effective-medium model with a tight-binding description of the shell energy has been shown to give a reliable description of the atomic structure of small *Cu* clusters. The calculated ground-state structures are in close agreement with experiment and other calculations and it reproduces the magic sizes observed experimentally. The coupling between the atomic and electronic degrees of freedom through **a** Jahn-Teller-like effect has been shown to have a number of important consequences. It **is** crucial for understanding the stability *of* the magic clusters *at* high temperatures and it is responsible for the strong evenodd oscillations which are experimentally observed. The atomic motion also plays a key role in the large broadening of the electronic levels observed in photoelectron spectroscopy. A phase transition of Cu_{13} at $T \simeq 450$ K is indicated by various analyses of the Monte Carlo results, whereas clusters of other sizes show more complex behaviour.

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